

RESEARCH ARTICLE

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Key Points:

- Ofu Island in Samoa has the highest $^3\text{He}/^4\text{He}$ ratio measured in the Pacific
- Major and trace element data for Ofu are presented here for the first time
- We present melt models evaluating the Ofu and Ta'u Island mantle sources

Supporting Information:

- ReadMe
- Appendix section III
- Figures A1–A2
- Tables A1–A3

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Ta'u and Ofu/Olosega volcanoes: The "Twin Sisters" of Samoa, their P, T, X melting regime, and global implications

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Abstract The Samoan islands of Ta'u and Ofu/Olosega (Ofu hereafter) are single shield volcanoes that have erupted alkali basalt for the past 70 and 440 kyr, respectively. They are 20 km apart, and are the easternmost subaerial expressions of the Samoan plume. The isotopic data for these islands are published; we report here the first major and trace element data for Ofu. The two islands are similar isotopically and in trace elements. Their high $^3\text{He}/^4\text{He}$ marks them as being a FOZO mantle end-member. By comparing data from both volcanoes, we test the efficacy of melting models in constraining the mantle compositions and their P-T of melting. We show that the mantle sources are similar, with spidergrams that peak at Ta ($4\times$ BSE), and Lu $\sim 0.5\times$ BSE. Melts and mantle sources both have Lu/Hf ratios that are too low to support measured $^{176}\text{Hf}/^{177}\text{Hf}$ ratios, pointing to a young enrichment event in the mantle source. Degrees of melting are 6.5% for Ta'u and 5.2% for Ofu. P-T of melting show a wide, overlapping range, but define a precise array; average values are 1475°C —77 km for Ta'u and 1550°C —110 km for Ofu. The deepest P-T estimate is 4.2 GPa and 1550°C . The P-T array is either a melting adiabat, or a mixing line of melts equilibrated at various depths. Kinetic modeling shows melt re-equilibration will be likely for ascent velocities (m/yr) less than 40/(conduit radius in cm)². P-T estimates for melting may typically be minimum values.

1. Introduction

Ta'u and Ofu are the two easternmost volcanic islands of the Samoa plume lineament, see Figure 1. The isotopic differences between the two volcanoes are very small, but distinct. This is in contrast to the very large isotopic variations displayed by all of the other volcanoes of the Samoa chain. In addition, while the chain as a whole has a strong EM2 character [Wright and White, 1987; Workman et al., 2004; Jackson et al., 2007b; Hart and Zindler, 1989], these two islands are devoid of EM2 material and have a FOZO signature, extending to very high $^3\text{He}/^4\text{He}$ [Jackson et al., 2007a]. This provides a unique opportunity to constrain the geochemical makeup of the FOZO mantle source, and to compare the pressure, temperature, and nature of the melting regime and melt transport system in these two Samoan volcanoes. Note that this paper presents the first major and trace element data for Ofu; sample localities are shown in Figure 1, and their location and description are given in Appendix Table A1. The major and trace element data for Ta'u, and all of the isotopic data considered here have been previously published.

First, we will briefly review the geology of these two volcanoes, to show that they are similar shield-stage alkali basalt volcanoes with a similar evolution, though they are at different geomorphic (erosional) stages. In section 3, the extant K-Ar age data are reviewed, showing that ages for Ofu are 440 kyr and younger, compared to ages of less than 70 kyr for Ta'u. Section 4 shows that Samoan volcanism occurs along two coeval trends separated by 50 km, with Pb isotopes becoming steadily higher along both trends as the active volcanism on the east ends is approached. Section 5 shows the resulting distinct but very small isotopic differences between Ta'u and Ofu. Section 6 compares trace element patterns, with both volcanoes being similar, and both having the distinct positive Nb, Ta, and Ti anomalies that are the hallmarks of high $^3\text{He}/^4\text{He}$ basalts and the FOZO mantle end-member [Jackson et al., 2007a, 2008]. Mantle source compositions are modeled in section 7 using the techniques of Minster and Allegre [1978], and Ta'u and Ofu give surprisingly comparable sources. Consideration of possible long-term isotopic evolution of these sources in section 8 suggests that these mantle compositions have undergone young modification. Section 9 uses the petrologic model of Lee et al. [2009] to derive P-T estimates for the melting of Ta'u and Ofu mantle. Kinetic modeling is used in section 10 to assess possible effects of melt-wall rock re-equilibration on these P-T estimates.

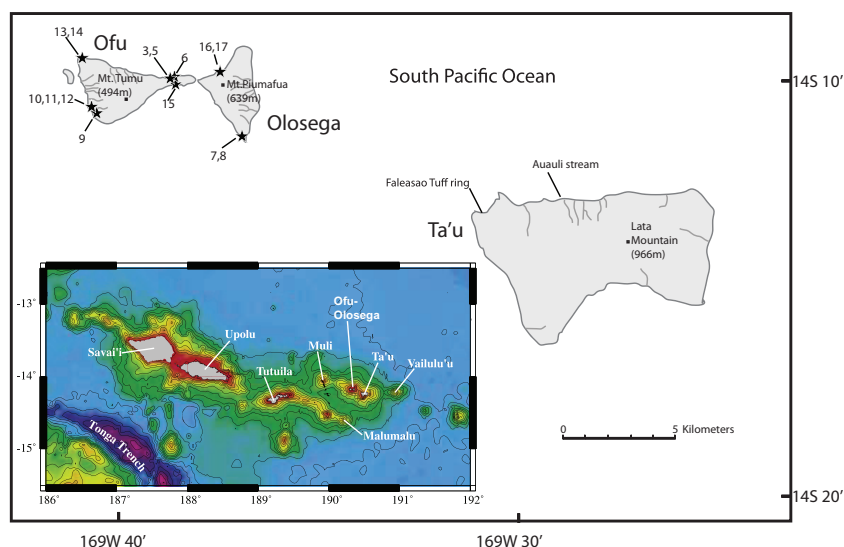


Figure 1. Location Map for Ta'u and Ofu-Olosega islands, and the Samoan Hot spot Chain. Ta'u is the easternmost island along the chain, with Vailulu'u submarine volcano being the active end of the chain. The locations of the Ofu samples (for which new major and trace element data are presented) are provided in the figure; for comparison with Jackson *et al.* [2007a], the sample ID's have the prefixes Ofu-04.

2. Overview of Geology and Field Setting

The first study of the geology of these islands was by Daly [1924], followed by Stearns [1944], and the latest is the comprehensive report by Stice and McCoy [1968], who defined five volcanic series on Ta'u and three on Ofol (Ofu and Olosega are parts of the same volcano). While these various volcanic series may be morphologically or stratigraphically meaningful, we have not found any geochemical signature that defines these series—each volcano is a single shield, decorated with minor postglacial tuff rings, cinder cones, and lava deltas. Furthermore, while Stice and McCoy assess the shape of these two volcanoes as derived from caldera formation, we agree with the earlier claim by Daly [1924] that the shape of Ta'u is created by major flank failure (foundering) on the south side. Our opinion is that a more extended history of multiple flank failures and debris avalanches has sculpted the Ofol volcano, from what earlier was a shape similar to present-day Ta'u, to its present shape divided into two adjoining islands. The exposure of the large headwall on the south side of Ta'u has been recently dated at 22 kyr [Williams *et al.*, 2014], suggesting its formation may be related to the low stand of sea level during the last glacial maximum. The deeper erosional level of Ofol volcano is also indicated by the presence of a major dike complex in the isthmus between Ofu and Olosega Islands (samples 04-02,-03,-05,-06).

Petrographically, both volcanoes are dominated by aphyric to mildly olivine or plagioclase phyric alkali basalt flows, with a few basanites, hawaiites, picrites, and ankaramitic dike rocks. All analyzed samples have SiO_2 less than 49%; the lowest Mg # is 35.7 (Table 1).

3. Age of Volcanism

K-Ar ages have recently become available for Ta'u, Ofu, and Olosega islands [McDougall, 2010]. Four samples from the oldest formation on Ta'u (Lata) range from 20 to 70 kyr (± 10 –20 kyr). Seven samples from the main shield formation on Ofu/Olosega (Tuafanua) range from 240 to 440 kyr (± 10 kyr). While these are total fusion K-Ar data, the rocks are reported to be unweathered, and there is no indication of excess Ar problems. In one case, adjacent flows on Ofu agree in age (240 and 250 kyr). In another case on Ofu, a lower flow gives a slightly younger age than an overlying flow (310 kyr versus 400 kyr). Two samples from the west rift of Ofu give $^{40}\text{Ar}/^{39}\text{Ar}$ plateau ages of 270 ± 140 kyr and 560 ± 70 kyr, while one sample from the south rift gives a plateau age of 490 ± 210 kyr [Koppers *et al.*, 2011]. These data provide compelling evidence that Ta'u is generally younger than Ofu/Olosega, as is also strongly endorsed by the relative morphologies of the two volcanoes.

Based on the Pacific plate velocity in this region of 71.3 mm/yr [Beavan *et al.*, 2002], and the spacing of the two volcanoes (\sim center of mass) of 20 km, we would expect a difference in age of \sim 280 kyr (for a stationary plume). This is remarkably close to the difference in average age of the samples reported by McDougall [2010] of 275 kyr. We are not inclined to present this as more than a general statement of the relative ages of these two volcanoes, as were we to make the same calculation for Vailulu'u (a submarine volcano that is the active locus of the plume) versus Ta'u (45 km separation), we would expect an age difference of 620 kyr, whereas the oldest dated sample from Ta'u is only 70 kyr (but clearly there are older basalts on Ta'u at greater depths in the shield). Evidence for young postglacial volcanism (<10 kyr) on Ta'u is given in Appendix III. Having set the "age" stage for Ofol and Ta'u, we can now address the issue of isotopic variations through time.

4. Temporal Isotope Variations

Although we will make the argument below that Ta'u and Ofu/Olosega are very similar geochemically, there are minor differences associated with an overall isotopic trend along the *en echelon* Vai and Malu trends in the Eastern Volcanic Province [Workman *et al.*, 2004]. This is illustrated in Figure 2, where $^{206}\text{Pb}/^{204}\text{Pb}$ is plotted against distance from the easternmost dredge from Vailulu'u. The overall shape of the two trends is remarkably similar, with a small dip in $^{206}\text{Pb}/^{204}\text{Pb}$ between 70 and 130 km. The Malu Trend is shifted \sim 20 km west of the Vai Trend, such that the currently youngest end of the Malu Trend (Malumalu seamount) would align with Ta'u Island in Figure 2, not the presently active Vailulu'u. Thus the Vai and Malu Trends are not only evolving coevally, but with very similar Pb isotopic signatures. The Vai Trend was initiated at \sim 0.97 Myr (Dredge 104 at 136 km) [Koppers *et al.*, 2011]. The Malu Trend initiates from eastern Tutuila, where the average age of the shield is \sim 1.3 Myr [McDougall, 1985]; the age of Dredge 108 at 123 km is 1.10 Myr [Koppers *et al.*, 2011]. The *en echelon* trends are not discernable prior to the generation of the Tutuila shield. Note that Vailulu'u is not exactly on the Vai Trend, but is shifted to the NE, as if to begin a new *echelon*. The distance between the two trends is \sim 50 km, and bears an uncanny resemblance to the Kea and Loa Trends on Hawaii, though these are somewhat closer together (30–40 km). These systematic Pb isotope variations will play a role in comparing isotopes between Ta'u and Ofol.

5. Isotopic Comparison of Ta'u and Ofu/Olosega

In general, there is limited isotopic variability within and between the islands of Ta'u, Ofu, and Olosega, with significant overlap. First, there are no significant isotopic differences between Ofu and Olosega Islands, so we will treat them as a single domain (as expected, as they are part of a single volcano, only separated morphologically by erosion). The clearest distinction between Ta'u and Ofol (Ofu-Olosega) can be made in $^{143}\text{Nd}/^{144}\text{Nd}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ space, see Figure 3. Except for two Ta'u samples that plot in the Ofol field (T54, T55), the two fields are distinct. These two Ta'u samples were collected very near to each other, from a flow running over a sea cliff, thus among the youngest samples analyzed from Ta'u. The two samples from the youngest formation on Ofu (Nu'u Fm) are on the lower edge of the Ofol field, but are not distinct (see Figure 3). One of the dike rocks on Ta'u (T44) is an isotopic outlier in several respects—it is the lowest $^{143}\text{Nd}/^{144}\text{Nd}$ and highest $^{87}\text{Sr}/^{86}\text{Sr}$ and Δ 208/204Pb of all of the samples studied here. This is in contrast to dike rocks from Ofol, which are abundant but not distinctive isotopically. In a statistical sense, the average $^{143}\text{Nd}/^{144}\text{Nd}$ for Ofol is only 41 ppm higher than Ta'u (0.512818 versus 0.512797), but this is meaningful at 2σ because the distributions for the two volcanoes are quite tight, with standard errors in each case of 7 ppm (1σ). The two volcanoes are only 44 ppm different in $^{87}\text{Sr}/^{86}\text{Sr}$ (0.704622 Ofol versus 0.704653 Ta'u), but this is not statistically significant. Figure 4 shows that the two volcanoes form arrays in $^{208}\text{Pb}/^{204}\text{Pb}$ - $^{206}\text{Pb}/^{204}\text{Pb}$ space that are colinear and overlap significantly. Finally, to provide some sense of the coherence between related geologic units, four samples were analyzed from across a coherent 50 m sequence of flows from the north shore of Ta'u; these samples are numbered in Figures 3 and 4, and show significant clustering. This establishes a rather typical feature of hot spot volcanism, where the isotopic variability increases with the scale length of the sampling [Hart and Zindler, 1989]; these two volcanoes also show much less isotopic variability than the Samoan chain as a whole, with a distinct absence of the characteristic Samoan EM2 component [Wright and White, 1987; Workman *et al.*, 2004; Jackson *et al.*, 2007b, 2010].

Table 1. Major, Trace Element and Isotopic Data for Ta'u, Ofu and Olsega Islands, Samoa (Hart and Jackson 2014)^a

	T10	T14	T16	T19	T21	T22	T23	T25	T27	T30	T32	T33	T38	T44	T45	T46	T47	T48
Location	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u	Ta'u
Rock Type	Basalt	Basalt	Basalt	Basalt	hawaiiite	Basalt	Basalt	picrite	Basalt	Basalt	Basalt	Basalt	Basalt	Ank'mite dike	Basalt	Basalt	Basalt	Basalt
Formation	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow	flow
Dredge Depth, m																		
Majors (wt%)																		
SiO ₂	46.60	46.25	46.71	47.14	48.61	46.51	47.49	45.23	48.12	46.41	48.02	47.32	45.34	46.84	47.93	46.56	48.31	46.13
TiO ₂	3.54	3.24	3.38	3.52	3.31	4.40	3.85	2.42	3.86	3.57	3.68	3.09	4.64	2.90	3.80	3.51	3.33	3.33
Al ₂ O ₃	12.41	10.84	11.27	12.12	17.66	14.47	13.44	8.37	14.17	11.80	13.50	10.61	13.93	10.63	12.81	12.19	11.51	11.75
FeO*	12.35	12.48	11.63	12.07	11.34	12.15	12.39	12.80	11.42	12.07	12.10	11.61	14.00	11.92	11.30	12.46	11.34	12.52
MnO	0.185	0.187	0.176	0.182	0.192	0.181	0.180	0.187	0.171	0.183	0.178	0.175	0.200	0.179	0.170	0.185	0.173	0.183
MgO	10.13	13.26	12.42	9.73	3.18	7.20	6.77	19.87	6.36	11.52	6.67	12.40	6.11	14.15	8.36	10.38	9.33	12.19
CaO	10.98	10.33	10.91	11.43	8.65	10.51	11.84	8.41	11.54	10.86	11.71	11.54	11.97	10.11	11.67	10.78	12.68	10.17
Na ₂ O	2.43	2.23	2.23	2.55	4.07	2.93	2.71	1.78	2.85	2.36	2.78	2.19	2.51	2.15	2.67	2.66	2.26	2.44
K ₂ O	0.920	0.820	0.860	0.860	2.070	1.150	0.920	0.660	1.060	0.830	0.930	0.749	0.930	0.790	0.870	0.850	0.710	0.870
P ₂ O ₅	0.449	0.371	0.389	0.390	0.929	0.516	0.419	0.282	0.457	0.408	0.422	0.334	0.388	0.331	0.429	0.427	0.358	0.414
LOI																		
Mg#	61.90	67.79	67.90	61.49	35.71	54.00	51.98	75.46	52.45	65.40	52.19	67.90	46.36	70.16	59.44	62.26	61.97	65.85
Traces (ppm)																		
Ni (xrf)	214	396	360	194	9	83	79	723	95	298	86	288	63	439	162	230	171	351
Cr	579	943	729	542	5	185	112	1406	290	682	100	988	59	764	549	515	654	687
V	336	333	340	345	182	369	373	254	334	313	356	308	489	283	353	323	346	308
Ga	19	20	20	20	27	22	22	13	25	19	22	18	24	17	25	20	21	22
Cu	72	116	32	110	101	70	105	69	131	102	133	61	120	92	99	115	106	85
Zn	111	126	102	127	150	124	123	111	122	121	121	103	121	115	113	130	112	120
Cs (icp)	0.120	0.120	0.090	0.090	0.580	0.190	0.050	0.162	0.090	0.170	0.100	0.106	0.080	0.090	0.140	0.200	0.160	0.050
Rb	19.9	18.0	19.8	18.8	49.1	26.3	16.4	15.9	21.4	19.0	19.0	15.7	16.7	17.1	18.4	19.2	14.7	17.8
Ba	178	164	177	164	406	248	190	125.8	208	166	180	140	203	140	170	176	160	171
Th	2.77	3.51	2.89	3.24	8.26	4.58	3.80	2.59	3.86	3.44	3.66	2.83	3.64	3.05	3.38	3.69	3.00	3.55
U	0.730	0.640	0.730	0.730	1.790	1.000	0.710	0.572	0.790	0.820	0.880	0.583	0.720	0.630	0.820	0.840	0.620	0.740
Nb	30.92	31.26	32.18	31.16	73.90	45.08	36.40	22.81	37.18	35.36	35.47	27.23	37.48	26.66	33.52	34.72	28.91	33.44
Ta	2.200	2.208	2.274	2.232	4.946	3.200	2.567	1.581	2.600	2.480	2.490	1.885	2.680	1.857	2.363	2.470	2.080	2.364
La	29.20	26.79	29.68	27.08	66.60	37.70	31.06	19.90	32.43	30.50	31.24	23.65	30.45	23.94	28.63	30.96	25.08	29.49
Ce	62.19	55.65	63.08	56.61	136.19	79.07	65.26	41.81	67.26	62.61	64.88	50.60	63.59	50.56	60.31	63.70	53.29	60.93
Pb	2.020		1.820	1.470	3.890	2.250	1.460	1.398	1.860	1.430	2.320	1.944	1.450	1.900	1.360	1.660	1.280	1.580
Pr	8.14	6.80	8.24	7.06	16.23	9.62	7.98	5.20	8.24	7.45	7.94	6.38	7.78	6.32	7.44	7.76	6.62	7.39
Nd	34.70	29.62	35.06	31.01	67.87	41.52	35.00	22.45	35.65	32.06	34.68	27.80	33.84	27.82	32.80	33.77	29.37	32.39
Sr	437	377	442	407	754	575	472	294	492	431	469	366	471	356	458	451	396	434
Zr	223	205	225	220	460	287	241	154	251	233	236	190	224	199	239	239	199	237
Hf	5.97	5.34	5.83	5.52	10.79	7.40	6.14	3.92	6.22	5.66	6.11	4.96	5.98	5.08	6.07	6.18	5.28	5.94
Sm	8.35	7.37	8.15	8.07	15.59	10.04	8.68	5.61	8.85	7.85	8.69	7.09	8.43	7.07	8.49	8.50	7.47	8.05
Eu	2.55	2.41	2.47	2.53	4.69	3.13	2.75	1.78	2.82	2.50	2.77	2.25	2.69	2.24	2.74	2.72	2.40	2.57
Gd	7.92	7.20	7.64	7.70	13.74	9.38	8.25	5.22	8.54	7.39	8.30	6.69	8.06	6.70	8.21	8.17	7.29	7.53
Tb	1.140	1.080	1.080	1.150	1.960	1.370	1.210	0.788	1.290	1.110	1.240	1.017	1.190	1.030	1.230	1.240	1.090	1.140
Dy	6.14	5.61	5.77	6.41	10.77	7.46	6.82	4.27	7.01	6.03	6.78	5.62	6.48	5.66	6.87	6.76	6.09	6.28
Ho	1.110	1.000	1.030	1.170	1.940	1.340	1.210	0.785	1.280	1.070	1.210	1.009	1.180	1.010	1.240	1.210	1.090	1.140
Y	29.41	25.74	27.91	28.94	50.43	34.20	30.48	20.21	32.55	27.50	31.25	26.16	29.69	25.28	31.49	31.09	27.53	28.85
Er	2.78	2.39	2.55	2.72	4.58	3.27	2.94	1.85	3.10	2.62	2.93	2.39	2.78	2.40	2.88	2.90	2.61	2.74
Tm	0.370	0.310	0.330	0.360	0.600	0.420	0.380	0.242	0.400	0.330	0.380	0.309	0.360	0.310	0.380	0.370	0.340	0.350
Yb	2.140	1.780	1.970	2.040	3.430	2.380	2.100	1.370	2.280	1.910	2.130	1.737	2.050	1.720	2.160	2.130	1.910	2.000
Lu	0.320	0.250	0.290	0.300	0.510	0.350	0.300	0.193	0.320	0.280	0.310	0.248	0.290	0.250	0.310	0.300	0.280	0.290
Sc	21.5	32.4	24.2	35.7	17.0	32.7	34.9	23.0	31.5	32.5	34.6	36.9	40.0	32.4	36.5	34.0	41.1	31.5
⁸⁷ Sr/ ⁸⁶ Sr	0.704657	0.704591	0.704605	0.704582	0.704751	0.704701	0.704706	0.704695	0.704561	0.704528	0.704588	0.704736	0.704651	0.705086	0.704434	0.704676	0.704504	0.704664
¹⁴³ Nd/ ¹⁴⁴ Nd	0.512789	0.512806	0.512818	0.512790	0.512796	0.512773	0.512796	0.512790	0.512806	0.512822	0.512797	0.512780	0.512785	0.512755	0.512816	0.512795	0.512800	0.512789
²⁰⁶ Pb/ ²⁰⁴ Pb	19.271		19.297	19.299	19.313	19.314	19.288	19.2266	19.304	19.305	19.208	19.284	19.328	19.246	19.337	19.231	19.253	19.249
²⁰⁷ Pb/ ²⁰⁴ Pb	15.602		15.601	15.600	15.594	15.606	15.614	15.6005	15.604	15.594	15.581	15.596	15.617	15.606	15.595	15.584	15.595	15.595
²⁰⁸ Pb/ ²⁰⁴ Pb	39.443		39.438	39.448	39.479	39.485	39.476	39.4024	39.500	39.382	39.332	39.445	39.532	39.549	39.425	39.369	39.366	39.397
³ He/ ⁴ He								13.26					16.62		14.86			
¹⁷⁶ Hf/ ¹⁷⁷ Hf														0.282948	0.282945			
Temperature, °C	1483	1500	1458	1468			1480			1476	1465		1563	1473	1431	1488		1498
Pressure, GPa	2.65	2.80	2.35	2.46			2.58			2.58	2.39		3.77	2.43	2.08	2.75		2.86

^aAnalytical data and methods for all of the Ta'u samples except T54 can be found in *Workman et al.* [2004]. Sr and Nd data were mostly by TIMS, the Pb was by ICP/MS, except T10, T16, T21, T55, and 74-1 were by TIMS. Isotope data for the Ofu/Olosega samples are from *Jackson et al.* [2007a]; Sr, Nd, and Pb were all analyzed on acid-leached powders by NEPTUNE ICP/MS. The Ofu/Olosega major and trace element data are presented here for the first time; analyses were provided by the GeoAnalytical Lab, Washington State University [Johnson et al., 1999; Knaack et al., 1994]. Analytical methods are described in *Hart and Blusztajn* [2006]. All Sr data are reported relative to 0.710240 for SRM987; precision is ± 15 –25 ppm (2s). All Nd data are reported relative to 0.511847 for the La Jolla Nd standard; precision is ± 15 –25 ppm (2s). All Pb data are reported relative to the SRM 981 values of *Todt et al.* [1996]; TIMS precision is $\pm 0.05\%$ amu; ICP/MS precision is ± 100 –200 ppm. Samples T16, T55, 04–09, and 04–15 were reanalyzed in July 2013 for ICP trace elements to assess the concordance of the Ta'u data analyzed in 2000, 2001, and 2002.

																	AVON	AVON	AVON
T51	T54	T55	04-03	04-05	04-06	04-09	04-10	04-11	04-12	04-13	04-14	04-15	04-07	04-08	04-16	04-17	74-1	74-4	75-10
Ta'u Basalt flow	Ta'u Basalt flow	Ta'u Basalt flow	Ofu Ank'mite dike	Ofu Ank'mite dike	Ofu Ank'mite dike	Ofu Ank'mite flow	Ofu Basalt flow	Ofu Hawaiiite flow	Ofu Picrobasalt flow	Ofu Hawaiiite flow	Ofu Ank'mite cumulate talus block	Ofu Ank'mite talus block	Olosega Basalt dike	Olosega Ank'mite dike	Olosega Basalt talus block	Olosega Gabbro talus block	Ta'u SE Rift picrite 2544 m	Ta'u SE Rift basalt 2544 m	Ta'u NE Rift basalt 2675 m
46.95	47.41	48.25	44.97	45.71	44.84	44.57	44.72	48.74	44.23	44.75	45.36	45.01	45.71	45.15	46.14	44.79	44.09	45.52	47.10
3.55	3.82	3.71	3.83	4.96	4.95	5.36	4.31	3.37	3.89	5.19	3.55	4.88	4.74	5.42	5.35	3.18	1.89	5.76	4.02
12.26	12.37	12.56	9.75	12.46	11.04	11.65	11.51	17.20	10.52	14.51	7.56	11.78	15.07	12.20	15.25	6.48	6.67	15.33	16.29
12.02	11.80	11.70	13.51	12.36	12.78	14.59	13.70	10.80	14.30	14.29	12.70	12.31	13.29	13.95	12.65	12.46	12.79	13.92	12.56
0.183	0.173	0.176	0.187	0.180	0.184	0.193	0.190	0.201	0.201	0.199	0.172	0.166	0.198	0.191	0.175	0.178	0.187	0.190	0.200
9.82	9.15	7.85	14.74	8.71	9.81	8.66	11.10	4.62	14.60	5.96	16.02	10.89	5.73	6.92	4.87	18.85	25.45	4.83	5.16
11.42	11.38	12.15	9.79	11.41	12.42	11.31	11.28	8.63	9.36	9.74	12.49	10.48	9.67	11.80	10.72	11.65	6.98	9.48	9.11
2.53	2.62	2.48	1.72	2.28	2.24	2.23	1.77	3.74	1.61	3.34	1.34	2.82	3.49	2.65	3.18	1.40	1.25	3.31	3.47
0.870	0.866	0.730	1.080	1.390	1.140	0.970	0.970	1.820	0.910	1.330	0.520	1.140	1.380	1.140	1.160	0.700	0.480	1.170	1.460
0.401	0.422	0.384	0.417	0.538	0.587	0.482	0.455	0.882	0.380	0.688	0.280	0.523	0.713	0.570	0.492	0.304	0.218	0.512	0.630
			2.01	2.57	1.37	−0.48	2.67	−0.44	0.83	−0.26	−0.95	0.65	−0.24	−0.16	0.05	0.14			
61.80	60.57	57.06	68.36	58.26	60.32	54.04	61.61	45.87	66.91	45.24	71.42	63.67	46.06	49.56	43.26	74.98	79.76	40.73	44.86
194	195	112	469	188	201	207	302	9	503	29	448	304	23	108	72	553	1019	36	5
540	653	331	836	397	533	276	826	6	1085	9	1437	604	5	192	58	1255	1655	32	15
341	322	355	295	355	364	407	352	226	317	424	328	326	382	398	358	281	187	472	282
22	23	22	19	22	21	23	21	30	18	30	15	23	31	24	24	15	13	24	26
110	109	104	76	95	80	127	75	5	77	15	48	74	11	111	118	55	69	145	21
126	100	124	125	127	132	142	128	171	147	171	110	122	166	143	131	110	103	142	140
0.100	0.116	0.100	0.233	0.254	0.330	0.161	0.306	0.396	0.175	0.309	0.041	0.282	0.258	0.341	0.213	0.164	0.130	0.150	0.390
17.5	15.1	12.0	27.3	34.7	25.6	21.1	22.4	42.7	20.7	30.7	11.4	24.8	31.4	26.4	27.6	19.3	10.5	22.3	27.6
177	172	147	196	271	249	216	242	388	214	279	112	230	301	232	238	141	87	237	264
3.43	3.34	2.65	3.26	4.24	3.96	3.25	3.65	6.60	3.29	5.02	1.95	3.89	5.24	4.06	3.54	2.18	1.51	3.98	4.94
0.670	0.704	0.600	0.760	0.995	0.991	0.807	0.831	1.534	0.688	1.232	0.445	0.969	1.296	0.987	0.893	0.525	0.420	1.110	1.100
32.16	32.90	27.93	34.85	46.01	47.80	37.99	39.30	68.82	32.57	55.52	21.52	45.80	57.20	46.74	42.23	25.96	18.94	45.78	51.26
2.274	2.271	1.981	2.573	3.389	3.444	2.771	2.814	4.772	2.358	3.921	1.591	3.264	4.088	3.397	3.053	1.864	1.320	3.380	3.59
27.81	27.15	23.86	29.6	38.5	37.9	32.8	32.9	58.3	33.7	44.8	18.6	34.6	47.8	38.1	31.4	20.3	15.66	37.13	43.08
58.49	57.81	51.28	62.9	81.9	80.5	69.2	68.4	122.3	59.9	94.8	39.6	72.8	101.0	81.0	66.3	43.1	33.14	77.39	90.50
2.050	2.396	1.950	2.095	2.671	2.608	2.432	2.252	4.120	2.028	2.898	1.164	2.412	3.039	2.599	2.350	1.355	1.060		3.260
7.28	7.21	6.52	7.81	10.16	9.99	8.69	8.53	14.94	8.15	11.77	5.16	9.02	12.55	10.16	8.27	5.39	4.30	9.44	11.08
31.76	32.49	29.68	34.0	44.1	44.1	38.0	37.7	64.7	35.9	51.8	23.4	39.7	55.1	44.9	36.2	24.1	18.16	40.30	47.98
427	459	427	390	546	599	503	495	929	331	723	292	601	767	573	639	310	217	660	617
223	237	214	218	288	293	250	248	442	214	351	148	280	362	297	253	151	113	259	327
5.72	6.08	5.60	5.89	7.56	7.56	6.72	6.56	5.78	8.77	4.14	6.98	9.22	7.76	6.39	4.12	3.01	6.79	8.31	
8.02	8.36	7.95	8.45	10.80	10.75	9.56	9.34	15.47	8.78	12.66	6.05	9.70	13.36	10.98	9.02	6.06	4.21	9.55	11.62
2.56	2.74	2.64	2.68	3.46	3.39	3.08	3.00	4.85	2.80	4.04	1.95	3.18	4.34	3.53	3.06	1.93	1.28	3.20	3.75
7.72	8.10	7.89	7.69	9.94	9.93	9.05	8.84	13.90	8.56	11.95	5.90	9.13	12.52	10.28	8.71	5.79	3.92	8.68	10.74
1.160	1.242	1.200	1.126	1.436	1.400	1.313	1.266	2.013	1.237	1.699	0.858	1.301	1.799	1.473	1.277	0.831	0.560	1.290	1.600
6.36	6.77	6.64	5.99	7.64	7.34	7.01	6.72	10.73	6.67	9.14	4.66	6.87	9.45	7.87	6.97	4.39	3.04	7.11	8.83
1.170	1.215	1.210	1.029	1.303	1.227	1.217	1.154	1.849	1.175	1.565	0.793	1.174	1.651	1.339	1.226	0.723	0.550	1.270	1.590
28.91	31.22	30.30	24.7	32.0	30.3	30.1	28.9	47.2	31.9	40.0	20.0	29.9	41.1	33.2	32.0	18.1	14.91	32.34	40.24
2.74	2.90	2.88	2.33	2.98	2.70	2.79	2.63	4.28	2.70	3.60	1.85	2.63	3.76	3.05	2.93	1.62	1.40	3.00	3.81
0.360	0.374	0.380	0.286	0.374	0.319	0.343	0.328	0.541	0.335	0.455	0.226	0.316	0.467	0.374	0.378	0.196	0.180	0.400	0.500
1.970	2.064	2.100	1.576	2.010	1.664	1.926	1.808	2.975	1.833	2.511	1.285	1.748	2.609	2.035	2.103	1.086	1.080	2.270	2.800
0.290	0.291	0.300	0.218	0.281	0.225	0.277	0.254	0.422	0.264	0.346	0.175	0.241	0.367	0.284	0.303	0.143	0.160	0.320	0.400
35.8	28	38.5	23.1	24.9	31.2	33.9	34.6	14.2	32.8	22.1	45.0	27.6	19.0	30.6	23.3	36.7	18.3	22.5	20.5
0.704614	0.704739	0.704545	0.704756	0.704793	0.704584	0.704538	0.704648	0.704635	0.704795	0.704625	0.704517	0.704559	0.704623	0.704703	0.704438	0.704498	0.704686	0.704815	0.704533
0.512794	0.512835	0.512815	0.512811	0.512814	0.512819	0.512835	0.512815	0.512804	0.512800	0.512816	0.512819	0.512822	0.512814	0.512802	0.512834	0.512844	0.512786	0.512784	0.512792
19.276	19.1775	19.178	19.257	19.254	19.189	19.165	19.187	19.184	19.202	19.176	19.126	19.141	19.177	19.207	19.216	19.160	19.290	19.314	19.266
15.591	15.5957	15.594	15.591	15.589	15.571	15.577	15.585	15.594	15.594	15.581	15.584	15.580	15.582	15.586	15.587	15.576	15.599	15.601	15.597
39.425	39.3369	39.298	39.418	39.412	39.202	39.208	39.261	39.299	39.403	39.264	39.164	39.169	39.268	39.291	39.276	39.173	39.443	39.477	39.39
	14.43		24.01	24.37	33.74	25.58	22.01		21.20		24.98	29.64		21.32	19.48	26.42	17.97		
			0.282956		0.282980						0.282971	0.282972			0.282974	0.282971		0.282933	
1468	1454	1446	1563	1490		1592	1559		1601			1497		1563					
2.48	2.32	2.10	3.65	2.90		4.20	3.64		4.24			3.16		3.85					

with the Ofu/Olosega data analyzed in 2005. Only Nb and Ta required minor corrections (5–10%), therefore all of the Nb-Ta data were recalculated to current calibration values; this and the duplicate analyses were used to correct the Nb-Ta data given in *Workman et al.* [2004] and used in this Table. Several Samoa samples (including T54 and 04-06) were also reanalyzed to high precision for Nd isotopes on the TRITON at DTM, and the agreement was better than 15 ppm [Jackson and Carlson,]. Hf isotope data are from *Salter et al.* [2011]. Mg number is calculated with FeO = 0.9 (total Fe as FeO). Note that the Mg# given in *Workman et al.* [2004] was calculated with FeO = 0.85FeO*. Pressure and Temperature data are given for selected samples, using the *Lee et al.* [2009] model; see text for details.

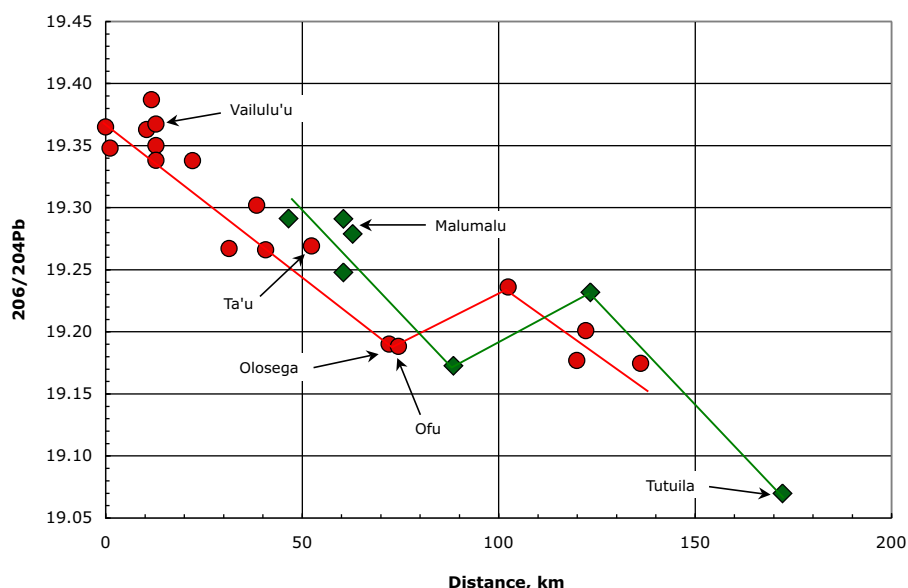


Figure 2. Variation of $^{206}\text{Pb}/^{204}\text{Pb}$ with distance along the Vai Trend (circles) and the Malu Trend (diamonds), Samoan Hot spot Chain. Each dredge or island location has been projected orthogonally onto a line of azimuth N64°W (the direction of Pacific Plate motion). Each symbol represents the average of data for samples from a single dredge (1–4 samples) or subaerial samples from a given island (Ta'u—20; Olosega—4; Ofu—10; Tutuila—11). Data from Workman *et al.* [2004]; Jackson *et al.* [2007a, 2010]; Wright and White [1987].

From Figure 2, it is clear that there is a gradual increase in $^{206}\text{Pb}/^{204}\text{Pb}$ toward the youngest end of the Samoan chain, and this slow “drift” in composition is likely causing the distinction between Ofol and Ta'u as seen in Figure 3. In the scheme of things, however, these isotopic differences are miniscule, and should not be taken as evidence for a significantly heterogeneous mantle.

6. Trace Element Comparison of Ta'u and Ofu/Olosega

Average trace element patterns (spidergrams) for the two islands are shown in Figure 5a, normalized to Primitive Upper Mantle (PUM). Appendix Figure A1 shows all of the individual spidergrams for the Ofol data set. The two average patterns are remarkably similar, with both showing a peak at Ta and depletions at K

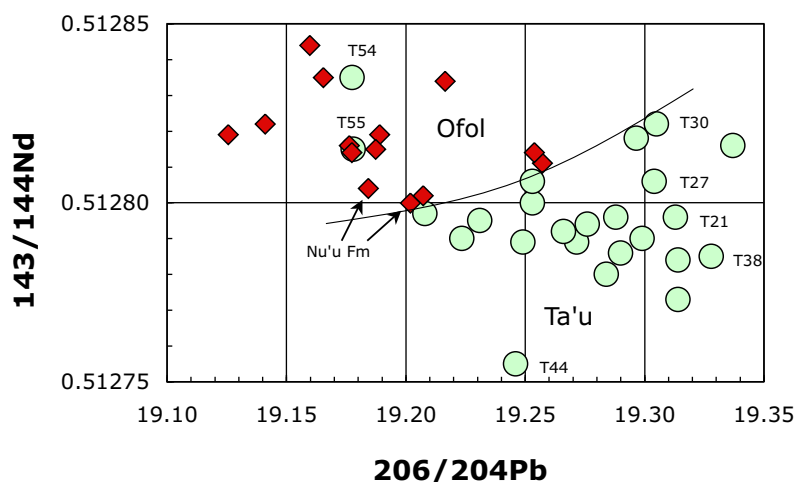


Figure 3. $^{143}\text{Nd}/^{144}\text{Nd}$ – $^{206}\text{Pb}/^{204}\text{Pb}$ plot showing the fairly distinct but minor isotopic separation between Ta'u and Ofol Islands. T44 is a dike sample from Ta'u. T54 and T55 are from a young posterosional flow on Ta'u, and are the only Ta'u samples that fall within the Ofol field. The other numbered Ta'u samples are from a 50 m flow sequence west of Auauli Stream, on the north flank of Ta'u (flows are ~2.5 m thick, with T21 at the bottom, overlain by T38, T30, and then T27 at the top). Two samples from the young posterosional Nu'u Formation on Ofu are kin to the other older Ofol samples. The Ta'u dredge samples 74-1, 74-4, and 75-10 are not distinctive on this plot.

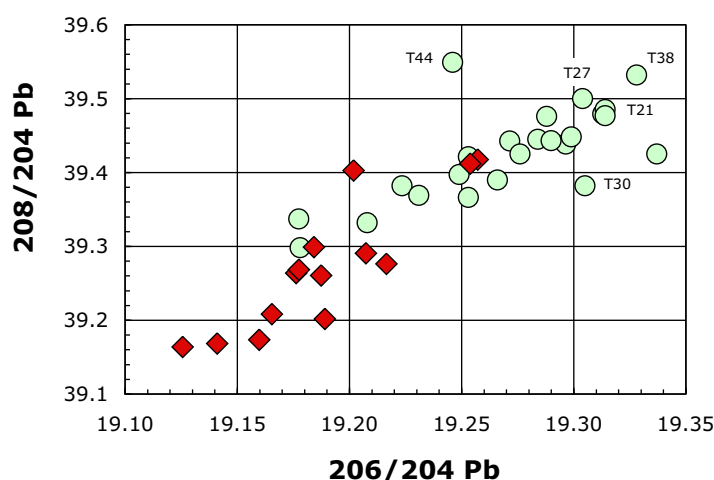


Figure 4. $^{208}\text{Pb}/^{204}\text{Pb}$ – $^{206}\text{Pb}/^{204}\text{Pb}$ plot showing the colinearity and partial overlap of Pb isotopes from Ta'u and Ofol Islands. T44 is a dike sample from Ta'u. The other numbered samples are from a 50 m flow sequence west of Auauili, on the north flank of Ta'u (flows are ~2.5 m thick, with T21 at the bottom, overlain by T38, T30, and then T27 at the top); these samples are distinctly clustered at high $^{206}\text{Pb}/^{204}\text{Pb}$ relative to the overall Ta'u array. The Ta'u dredge samples 74-1, 74-4, and 75-10 are not distinctive on this plot.

and Pb, as is typically observed in OIB from HIMU, EM1, or EM2 hot spots [Hart and Gaetani, 2006]. There is a small Ba depletion as is typical only in EM2 [Workman *et al.*, 2004]. There is a small Ti peak, but no significant Eu anomaly ($\text{Eu}/\text{Eu}^* = 1.01$ and 1.00). The average Ofol pattern is more enriched than average Ta'u for most of the elements, with the patterns crossing only at Er. It is tempting to attribute the enrichment in the Ofol samples to their being generally more evolved (loss of olivine and/or clinopyroxene), or the Ta'u samples being more influenced by accumulation of olivine and/or clinopyroxene. The average Mg# is lower for Ofol than for Ta'u (57.8 versus 59.4), but this small difference cannot account for the observed difference in the spidergrams. Furthermore, the average MgO and CaO of the two islands are identical (10.11 versus 10.10, 10.77 versus 10.69), so differences due to olivine or clinopyroxene control during magmatic differentiation can be ruled out. It is likely then to be a source effect, whereby either the Ofol source is more enriched than the Ta'u source, or the degree of melting is smaller. We show below that it is most likely a degree of melting effect.

To pursue this question, in Figure 5b we show the difference between the two islands as a ratio of Ofol/Ta'u, in percent. Ignoring the high Cs and Rb values for Ofol, that may be alteration related, we see a general enrichment level in Ofol of about 20% for elements from U to Gd, with peaks above this level of 5–10% at Ta, Pb, and Ti. The heavy REE from Gd to Lu show a very linear negative slope, reaching a depletion at Lu of ~10%. The REE “difference” pattern is thus composed of two linear segments, one horizontal from La to Gd, and the negative sloped segment from Gd to Lu. This pattern is not difficult to generate as a melting effect, using variations in degree of melting, and modal source differences in clinopyroxene and garnet.

Jackson *et al.* [2008] showed that all high $^3\text{He}/^4\text{He}$ ocean island basalts (OIB) are characterized by positive spidergram anomalies in Nb, Ta, and Ti. Therefore, the small Ta “peak” shown in Figure 5b is examined further by noting that the Nb-Ta correlations are excellent, with a minor (2.5%) distinction between Ofol and Ta'u: Nb/Ta for Ta'u = 14.22 ± 0.10 , and for Ofol = 13.87 ± 0.13 (2σ standard errors). These values straddle the BSE (bulk silicate earth) value of 14.0 ± 0.3 proposed by Munker *et al.* [2003]; both islands also have Nb/Ta that is very similar to N-MORB (14.0 ± 0.4) [Munker *et al.*, 2003]. If real, the small difference between Ta'u and Ofol could be created by partitioning effects of a number of minor minerals that may be present in the mantle source, e.g., carbonate, rutile, biotite/phengite [Dasgupta *et al.*, 2009; Stepanov and Hermann, 2013; Schmidt *et al.*, 2004]. It may be noted that Pfander *et al.* [2007] determined an average Nb/Ta = 15.9 ± 0.6 for a large suite of samples from six end-member OIB hot spots; however, out of 48 analyzed samples, none were as low in Nb/Ta as the Ta'u and Ofol values. Notably, the lowest Nb/Ta ratios reported by Pfander *et al.* [2007] were on samples from Samoa. Not only does this analysis confirm the broad correlation between high $^3\text{He}/^4\text{He}$ and positive Nb, Ta, Ti anomalies demonstrated by Jackson *et al.* [2008], but it draws the mantle end-member FOZO into the correlation since high $^3\text{He}/^4\text{He}$ was noted as a main signature of FOZO [Hart *et al.*, 1992; Jackson *et al.*, 2007a].

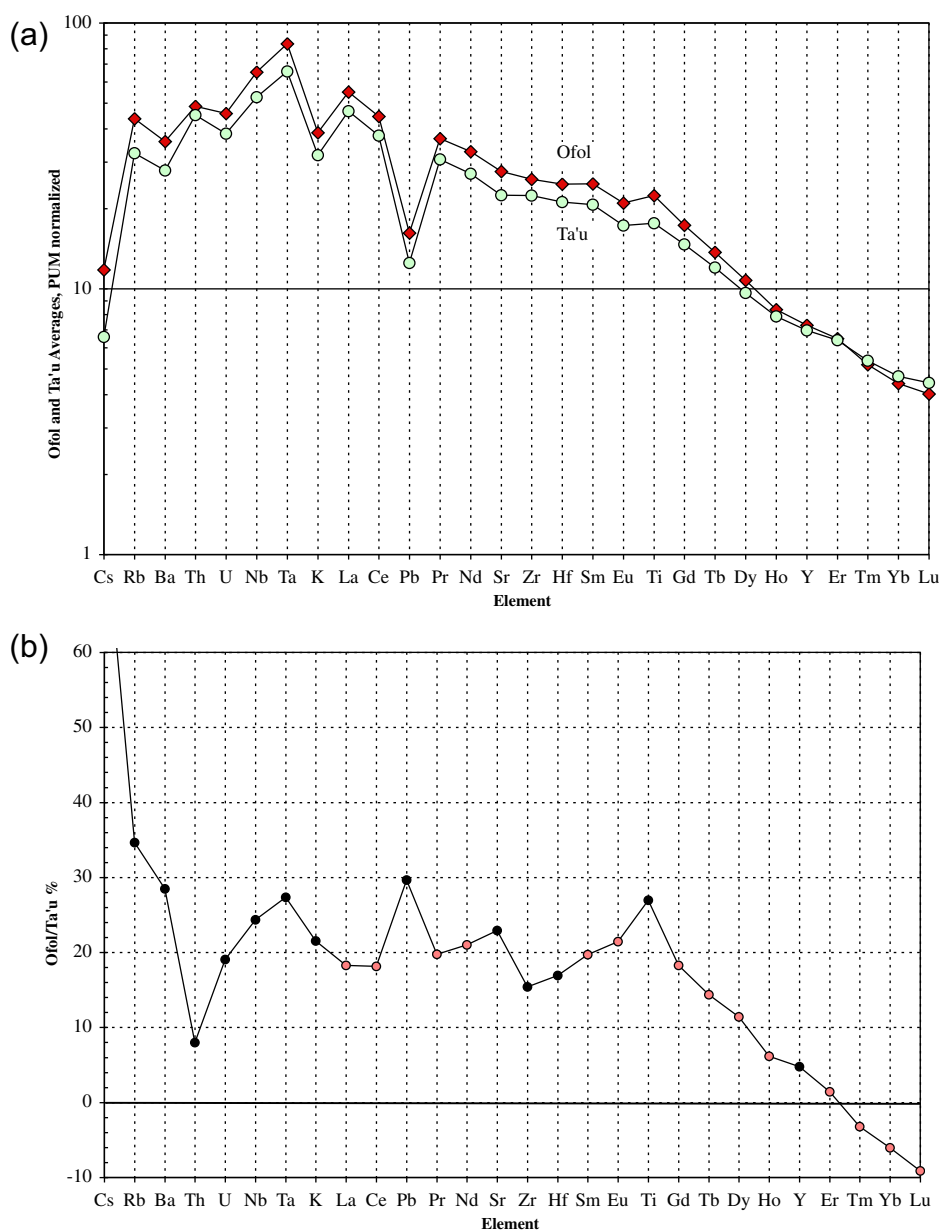


Figure 5. Trace element spidergrams for the average of all 14 samples from Ofof and all 24 samples from Ta'u. The dredge samples are not included. (a) The average spidergram for each island, normalized to primitive upper mantle (PUM [McDonough and Sun, 1995]). (b) Comparison of average elemental differences, in %, between the two islands (REE symbols are in orange for clarity). Note that the McDonough and Sun normalizations are based on Nb/Ta = 17.8, whereas more recent considerations argue for a lower value, = 14.0 ± 0.3 [Munker et al., 2003]. This would decrease the Nb-Ta slope on the spidergrams in Figure 1a, but would not change the Ta "peak" in Figure 1b.

7. Mantle Source Spidergrams

By use of linear regression of reciprocal melt concentration plots (e.g., $1/\text{Sm}$ versus $1/\text{Th}$), we can gain information about both elemental concentration ratios in the mantle source, and bulk partition coefficients during melting [Minster and Allegre, 1978; Sims and DePaolo, 1997; Hart et al., 1997]. For incompatible elements, or elements of similar bulk partition coefficient, the slope on such a plot is equal to the concentration ratio in the mantle source (e.g., Th_0/Sm_0). The intercept is a function of the bulk partition coefficients and element concentration in the mantle source (e.g., Intercept = $(D_{\text{Sm}} - D_{\text{Th}})/\text{Sm}_0$, for small D_{Th}). This linearization is functional for incompatible elements, and for modal or nonmodal batch melting. While not strictly linear for the more likely melting process involving aggregated fractional melts, the nonlinearity is of minor

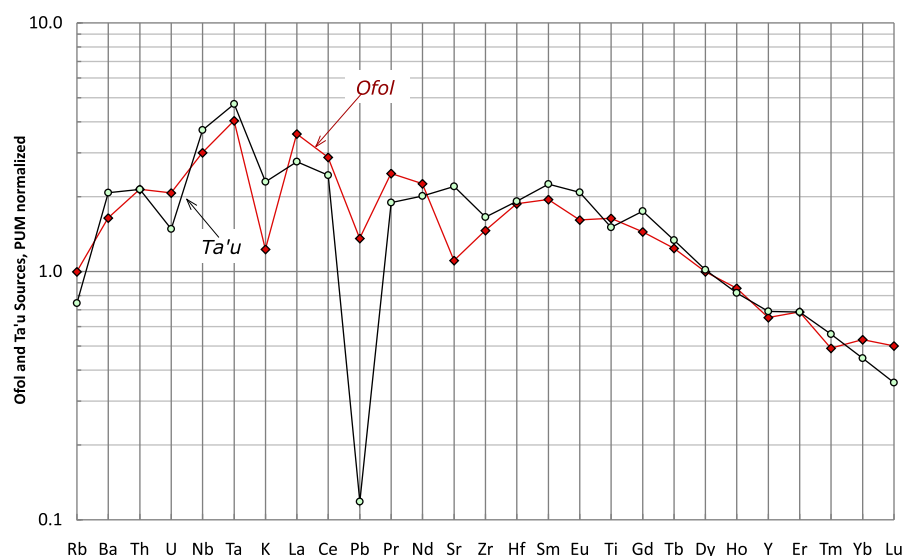


Figure 6. Mantle source spidergrams for Ta'u and OfoI (see text). Only the samples filtered for Cpx control or olivine accumulation are used here; these samples are also corrected to a common Fo 87 olivine equilibrium. Sample 04–12 from the posterosional Nu'u Formation was excluded from the regressions because it appears anomalous on most element–element plots. The regressions that were used to generate these source compositions are given in Appendix Table A2. The mantle composition data are given in Appendix Table A3.

consequence for highly incompatible elements and small degrees of melting [Hart *et al.*, 1997]. Effects due to differentiation during magma ascent must first be assessed, since some samples have clearly accumulated olivine, and some have differentiated to Cpx crystallization. For our modeling, we have limited the data set, based on CaO–MgO relationships, to samples nominally free of Cpx control or olivine accumulation. We only used samples that fell within an area in CaO–MgO space bounded by lines connecting the CaO–MgO vertices = 12.0%–5.0%, 9.0–15.0, 10.2–15.0, 13.3–5.0. In addition, the trace element contents of this selected set of samples have then been adjusted for the olivine masses that needed to be added or subtracted to reach a common Fo 87 equilibrium olivine (Fo 87 is the most magnesian olivine phenocryst in any of the lavas). This set of 15 Ta'u and 7 OfoI samples will be used for all further trace and major element modeling.

Using the filtered data set and pair-wise regressions of $1/C_A - 1/C_B$ data, with iteration where necessary, we have derived full spidergrams for both the Ta'u and OfoI mantle sources—see Figure 6. (A listing of all the regressions is given in Appendix Table A2. Appendix Table A3 lists the final concentrations for the two sources, and the partition coefficients for the OfoI data set.) These sources have been “pinned” at Er as this is where the basalt spidergrams cross (Figure 5b). The particular Er concentration is a free parameter, but was chosen at 0.685 PUM (0.300 ppm) based on actual data from peridotites in the McDonough (1990) database, and the discussion in Hart *et al.* [1997]. Overall, the Ta'u and OfoI sources show remarkably similar spidergrams across the whole range of incompatibility, with a few exceptions. The heavy REE (Tm, Yb, Lu) show a curious flattening for OfoI relative to Ta'u, significant at $\sim 1.5\sigma$. Both sources exhibit the marked K and Pb depletions that characterize most oceanic basalts [Hart and Gaetani, 2006], showing that these depletions are present in the source and are not produced by the melting process. The source for the HIMU suite from the Hobbs Lineament, Antarctica, showed similar K depletions [Hart *et al.*, 1997]. The difference between the Ta'u and OfoI K anomalies is significant at 2σ , the difference between the Pb anomalies is not demonstrable due to a very uncertain value for Pb in Ta'u. Sr shows a small positive anomaly in the Ta'u source, and a marked negative anomaly in the OfoI source; the difference is significant at $\sim 1.6\sigma$. Neither the Ta'u Ti anomaly nor the OfoI Th anomaly is provable at the 1σ level. However, the Nb/U ratio is significantly higher in Ta'u than in OfoI by $>2\sigma$ (79 versus 47).

If OfoI is a FOZO mantle end-member, as argued based on its high $^3\text{He}/^4\text{He}$, then Figure 6 is the first estimate for the geochemical makeup of FOZO. It is certainly not a “primitive” or “primordial” source as might be argued based on the high $^3\text{He}/^4\text{He}$. It is also not “depleted,” as was argued by Hart *et al.* [1992]. The LREE are more than 10 times enriched over the depleted MORB mantle of Workman and Hart [2005]. The LREE

are also ~ 3 times more enriched than the HIMU mantle defined by Hart *et al.* [1997], and the overall REE pattern is slightly convex upward, versus the markedly bowl-shaped REE pattern for the HIMU source. But, as we will discuss in the next section, this FOZO mantle composition does not appear to have a long-time integrated history.

From the source spidergrams (Figure 6), the olivine-corrected rock compositions, and the bulk partition coefficients (D , given in Appendix Table A3), it is straightforward to calculate the fractional degree of melting (F) for each sample and each element (only the Ofol data provided intercept estimates of reasonable precision to constrain bulk D 's. These D 's were thus also used for the Ta'u calculations, given the similarity of the source spidergrams, and the insensitivity of F to D , when D is small). Using the 15 most incompatible elements, from Ba to Ti (excluding Rb, K, Pb, and Sr), there is typically good agreement for F in a given sample (1σ SE typically ~ 0.005 for Ta'u samples and 0.001 for Ofol samples). F for the 14 Ta'u samples ranges from 0.058 to 0.076 , averaging 0.0650 ± 0.0026 (2σ standard error); for the seven Ofol samples, the range is 0.044 – 0.060 , with a mean of 0.0516 ± 0.0045 . Thus there is only overlap in F with one sample (from Ofol), and the means are strongly different at 2σ , with F for Ta'u being 26% greater than for Ofol. This follows directly from the 20 to 30% trace element enrichment in Ofol lavas compared to Ta'u, as shown in Figures 5a and 5b. For comparison, the degrees of melting for these FOZO sources may be compared to that determined for a HIMU source (Hobbs Lineament, Antarctica [Hart *et al.*, 1997]), using the same techniques: average F for Coleman Nunatak lavas = 1.6% versus 5.2% for the Ofol mantle source. This difference is consistent with the fact that the Hobbs volcanoes are small, erupted through continental lithosphere, and are rift related, not plume-driven.

This is not to imply that melting under Ofol was fundamentally different from that under Ta'u. There are simply more samples from Ofol that were derived from lower F melting domains than from Ta'u. Interestingly, the P-T modeling discussed below shows that these Ofol samples also have a higher mean P and T than Ta'u, consistent with more rapid melt migration from greater depths in the Ofol plumbing system. In other words, Ofol melts were effectively removed from the system at deeper levels and lower degrees of melting. Under Ta'u, melting and ongoing equilibration between mantle and melt continued during upwelling to shallower depths than under Ofol, leading to the overall higher degrees of melting.

8. Mantle Source Evolution

In common with many other OIB suites of HIMU/FOZO taxonomy, the Ta'u-Ofol volcanic suite shows a disconnect between low (or subchondritic) Lu/Hf ratios and high (or superchondritic) Hf isotopes. This is also true for Sm/Nd and Nd isotopes for Ofol lavas (Ta'u lavas are slightly superchondritic in Sm/Nd). While recognized for decades, this dilemma is commonly dismissed by asserting that the melting process has simply decreased the parent/daughter ratios relative to the mantle source. To the extent that the source compositions we have modeled above (e.g., Figure 6) are valid, we show that both the source and the melts are LIL-element enriched, or subchondritic in Lu/Hf. So sources with subchondritic Sm/Nd and Lu/Hf have produced superchondritic ratios of $^{143}\text{Nd}/^{144}\text{Nd}$ and $^{176}\text{Hf}/^{177}\text{Hf}$.

For the Sm/Nd system, the Ofol source has precise enough Sm/Nd data to show this effect: the weight ratio is 0.280 ± 0.009 , whereas the $^{143}\text{Nd}/^{144}\text{Nd}$ ratio is 0.512814 (average of the seven basalts used to define the source). Chondritic values are: 0.324 and 0.512630 . The effect is much more clearly shown for Lu/Hf, where even a visual inspection of Figure 6 shows very low subchondritic Lu/Hf ratios for both Ta'u and Ofol.

For Ta'u, the Lu/Hf (weight ratio) is 0.0405 ; for Ofol it is 0.0631 ; both are markedly lower than the chondritic ratio of 0.2367 . Ta'u and Ofol have very similar $^{176}\text{Hf}/^{177}\text{Hf}$ isotope ratios (0.282946 and 0.282964 [Salters *et al.*, 2011]), both clearly higher than the chondritic value of 0.282785 (definitive chondritic values for the Sm/Nd and Lu/Hf systems are from Bouvier *et al.* [2008]).

Extrapolating back in time on an isotope evolution plot, the Lu/Hf evolution trajectories for Ta'u and Ofol do not converge with the chondritic evolution curve. They do intersect the mantle evolution curve for depleted MORB (DMM, from Workman and Hart [2005]), at about 300 Myr. They would likewise intersect an evolution curve for the Early Depleted Reservoir (EDR) formed by early segregation of an enriched, now hidden, reservoir, proposed by Boyet and Carlson [2005, 2006] to solve the problem of the superchondritic $^{142}\text{Nd}/^{144}\text{Nd}$ ratios of the accessible earth. Using the average Lu/Hf system values for the EDR as given by

Jackson and Jellinek [2013] would lead to an intersection with Ta'u/Ofol of about 300 Myr, the same as with DMM. In either event, a young process of trace element enrichment of a long-term depleted source must have taken place. This enrichment process may be metasomatic in nature and separate in time and space from the melting process producing Ta'u and Ofol volcanism [e.g., Frey *et al.*, 1978; Zindler *et al.*, 1979; Roden *et al.*, 1984a], or it may be an integral part of the melting process (termed "consequent metasomatism" by Roden *et al.* [1984b], and essentially an autometasomatic process). There is abundant evidence for metasomatism in mantle xenoliths from Samoa, with isotopic character similar to erupted melts [Hauri *et al.*, 1993; Hauri and Hart, 1994]. While most easily considered as an autometasomatic signature created by the infiltration and passage of Samoan melts, we cannot rule out an earlier unrelated metasomatic event that in effect created the present Samoan mantle source. For example, underplating and lithospheric metasomatism from earlier hot spot tracks in the Samoa region may be an important process [Jackson *et al.*, 2010]. And finally, a young metasomatic process could generate the minor differences in the Ta'u and Ofol source compositions (Figure 6), without concomitant significant differences in isotopic character. Thus, while we have defined a mantle source composition for these two volcanoes, it is not the same as the long-term source in which the isotopes were incubated.

9. P-T Conditions of Melting

To further compare the melting regimes and P-T conditions between Ta'u and Ofol, we have used the filtered data set described above (section 7) to calculate P-T of melting, using the parameterized model of Lee *et al.* [2009]. While models such as these are not yet fully realized, due to inadequate phase equilibria data, the similarity of the Ta'u and Ofol mantle sources may lead to a better intercomparison than the absolute precision expected from the model implementation. For example, the major free parameter in the model is the selection of the Fo content of the mantle source. A variation of only 1 Fo unit creates a P and T shift of ~ 0.5 GPa and $\sim 50^\circ\text{C}$ (at Fo89). But, given the isotopic and trace element similarity of Ta'u and Ofol, we can safely assume the same Fo for both sources.

Figure 7a shows the P-T data calculated for the filtered Ta'u and Ofol data sets (for Fo89 mantle olivine), along with a 1550°C potential temperature adiabat, and both a dry and a damp solidus. The Ta'u and Ofol data fall closely along a single trend, which we for now identify as an isentropic melting adiabat. This melting adiabat falls ~ 50 – 70°C above a dry peridotite solidus, and 200 – 240°C above the most likely damp solidus (see figure legend for how this was estimated). It is unlikely that this melting adiabat is following a depleted solidus, as Wasylenki *et al.* [2003] showed that the solidus for a very depleted peridotite was only 55°C higher than that for a fertile (PUM or BSE) peridotite [see also Robinson *et al.*, 1998]. Workman and Hart [2005] showed that the depleted upper mantle (DMM) is only mildly depleted relative to PUM, and the Samoa source mantle is probably not as depleted as DMM. Nevertheless, the tight array shown by the Ta'u-Ofol data (Figure 7a) suggests a strong control by phase equilibria, probably a melting reaction boundary such as $\text{Ol} + \text{Cpx} + \text{Gnt} = \text{Opx} + \text{Liquid}$ [e.g., Kinzler and Grove, 1992; Longhi, 2002; Walter, 1998].

Plots such as Figure 7a are typically interpreted as indicating a range of actual potential temperatures, leading to averaging as a means of comparing potential temperatures of volcanism within or between various geotectonic environments. There is also the recognition that adiabatic cooling can occur during melt transport, thus arguing for the use of the maximum potential temperature for a given suite of volcanics [Herzberg *et al.*, 2007; Rooney *et al.*, 2012; Reid *et al.*, 2012; Putirka *et al.*, 2007; Gazel *et al.*, 2012; Courtier *et al.*, 2007]. Kelley *et al.* [2010] offer a more cautious interpretation of such P-T data, suggesting they may represent "final conditions of equilibration of a single batch melt," or a "continuously interacting melt that percolates upward through the solid mantle and reacts continuously along its path." We will evaluate this proposition below, using a formal diffusion closure-temperature approach; we find it quite likely that re-equilibration of melts during ascent will be the norm.

The range of potential temperatures captured by Figure 7a is from 1400° to 1550°C . Thus the shallowest depth of origin or of re-equilibration is ~ 60 km, which is within the lithosphere under Samoa, unless substantial lithospheric thinning has occurred as a result of erosion by the plume. Backtracking of the highest P-T sample along a curved melting adiabat will clearly lead to a much higher potential temperature at the point where the melting adiabat reaches the damp solidus. Thus the maximum potential temperature suggested by this data could be of order 1600°C , and the depth of intersection with the chosen damp solidus

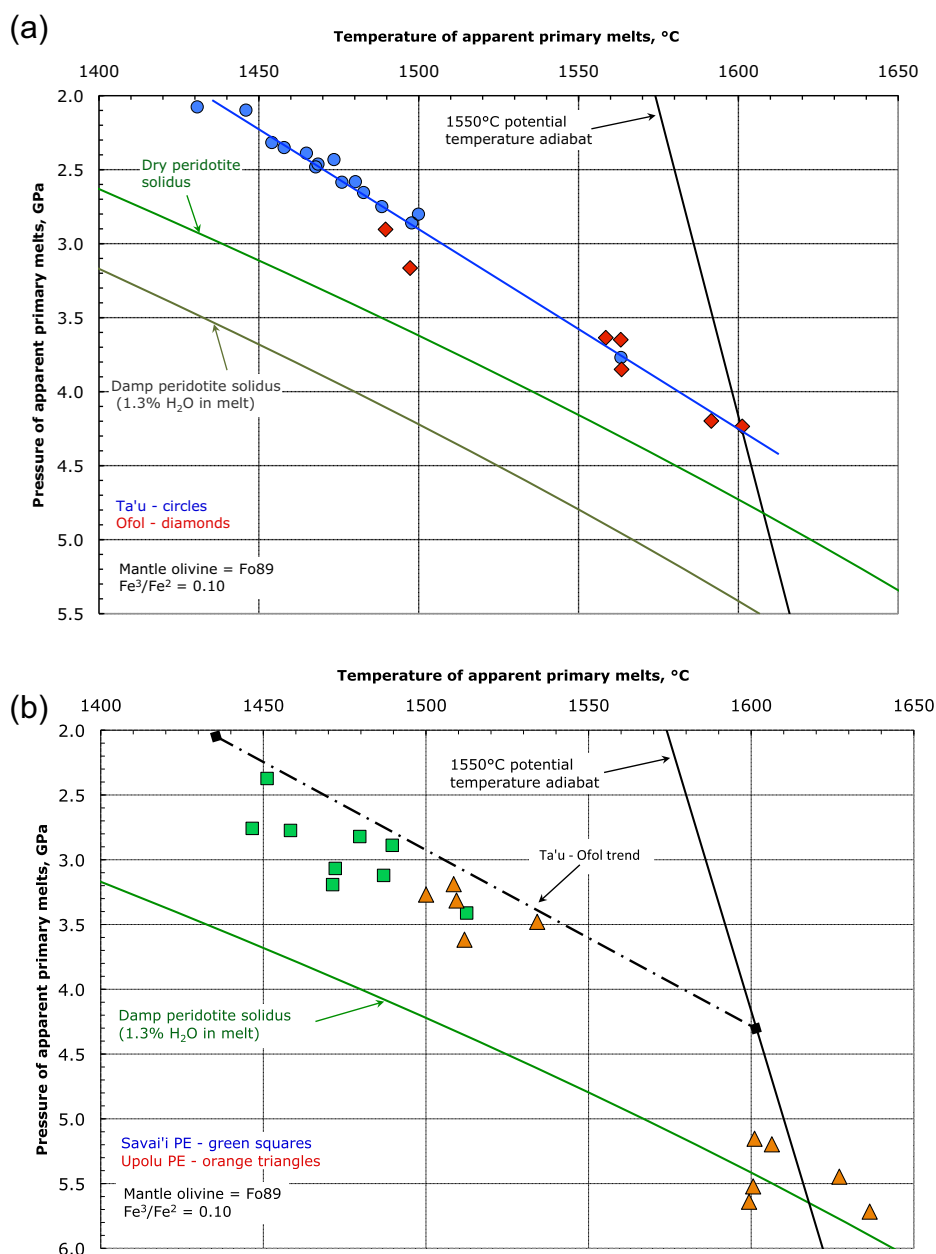


Figure 7. (a) Pressure-Temperature estimates for 15 basalts from Ta'u Island (circles), and seven basalts from Ofu-Olosega Islands (diamonds), selected to be unaffected by Cpx differentiation (see text). A mantle olivine composition of Fo89 was assumed for both suites, using the model of Lee *et al.* [2009]. This choice is based on the assumption that the Ta'u-Ofu mantle will not be more depleted than the upper depleted MORB mantle, and this mantle is fairly well constrained to be ~Fo89.5 [Workman and Hart, 2005]. The blue line is a best fit to the data from both Ta'u and Ofu Islands ($p = 0.0134 \cdot T - 17.22$). Also shown (green) is an anhydrous peridotite solidus [Katz *et al.*, 2003], and a "damp" peridotite solidus (dark green) using the parameterization of Katz *et al.* [2003], refined with the new data of Tenner *et al.* [2012]. The measured water contents of 25 glasses from two dredges near Ta'u Island were used (average water content = 1.30 wt %, shown to be unaffected by degassing [Workman *et al.*, 2006]). This corresponds roughly to a subsolidus mantle water content of ~150 ppm [Workman *et al.*, 2006]. The solid mantle adiabat shown in black is for a potential temperature of 1550°C (slope = 0.4°C/km). (b) Comparison of the Tau-Ofu P-T trend from Figure 7a with post-Fagaloa lavas from Upolu (11 samples) and Savai'i (nine samples). Fagaloa is the oldest mapped unit (shield) on Upolu and Savai'i [Kear and Wood, 1959]. These samples were selected from the data in Workman *et al.* [2004] solely based on MgO > 11%; the data were quite clustered in CaO-MgO space, with no evidence for Cpx control. Samples include all of the post-Fagaloa series on Upolu (diamonds) and Savai'i (circles). These series are, in decreasing age: Salani, Mulifanua, Lefaga, Puapua, and Aopo [Kear and Wood, 1959]. The high-pressure group includes only samples from Upolu (Mulifanua, Lefaga and Puapua series). (c) Comparison of the Tau-Ofu P-T trend from Figure 7a (diamonds) with the 26 highest MgO N-MORB glasses from Melson *et al.* [2002]. These glasses have MgO from 8.83 to 9.80%, are from a water depth range of 2500–3625 m in active near-ridge environments only (ages <6 Myr). Two samples with anomalously high Al₂O₃ were not used, nor were one plag-phyric glass and one sample from the Red Sea. This group of N-MORB glasses defines a very tight trend, subparallel to, but ~120°C above, the anhydrous solidus; the maximum potential temperature is ~1365°C. Also shown here is a melt mixing line between an average of the six lowest temperature primary N-MORBs and an average of the five highest temperatures Ofu-Ta'u samples; tic marks are at intervals of 5%.

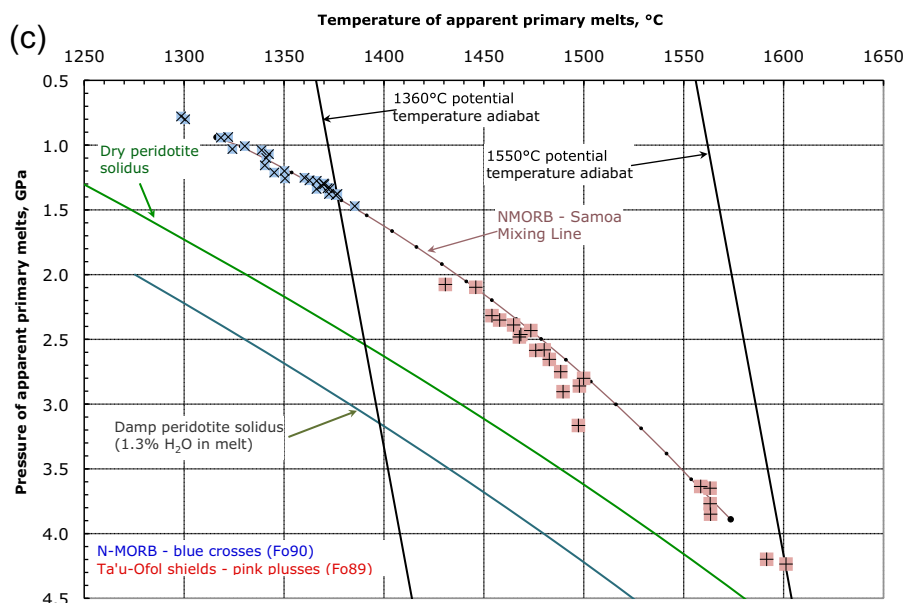


Figure 7. Continued.

of order 150–200 km. Were we to use a higher mantle Fo content for the modeling, the P-T of the high temperature end of the array would shift to even higher P and T. For example, *Hauri and Hart* [1994] report olivine compositions from a suite of xenoliths from Savai'i, and there is a clustering of 13 samples that averages Fo91.1. While these are harzburgites that are possibly derived from the sub-Samoan lithosphere, they have a strong EM2 Samoan isotope signature. Using Fo91.1 instead of Fo89 for the Ta'u-Ofol suite leads to a P-T for the highest pressure primary magma of 1740°C—6.94 GPa; this represents a temperature increase of 140°C, and a “depth” increase of 80 km.

The western islands of Upolu and Savai'i have entered a posterosional or rejuvenated stage of volcanism, and it is of interest to compare the P-T conditions of this stage with the shield stages on Ofol and Ta'u discussed above; Figure 7b makes this comparison. All of the Upolu-Savai'i data lie below the Ofol-Ta'u trend, by up to 1 GPa. While there is partial overlap of Upolu with Savai'i at shallow depths, there is a high P-T cluster of Upolu data that falls around the 1550°C potential temperature adiabat that also marks the high temperature end of the Ofol data. This Upolu cluster falls off a linear extension of the melting adiabat, but may mark the “turning” of this adiabat as is required for it to reach the damp solidus. In any event, this Upolu cluster clearly supports a potential temperature adiabat of >1570°C, as argued above for the Ofol data.

To add context to the Ta'u-Ofol P-T data, we have selected a set of high MgO N-MORB glasses from the Smithsonian glass data set [Melson *et al.*, 2002], and used the Lee *et al.* [2009] model to calculate P-T conditions for these primary N-MORB melts; see Figure 7c. These data define a very tight melting array, shifted far down in P and T from the Ta'u-Ofol array. The slope of the N-MORB array is slightly less than that of the Ta'u-Ofol array, but plausibly could represent a slightly curved extension of that array. This argues for the generalized phase equilibria control of the P-T melting trajectories as discussed above. The N-MORB melting adiabat falls above the dry solidus by ~120°C, but is subparallel to it, again demonstrating that the melting adiabat is not controlled by the solidus (of course, all of this assumes the P-T parameterization of Lee *et al.* [2009] is accurate). The range of potential temperatures represented by the N-MORB data is from 1290° to 1360°C, with a depth range from 25 to 45 km. Remember, this is not the depth of origin, but the depth of last equilibration with residual peridotite. Backtracking the melting adiabat to the dry solidus would put the maximum depth of melt origin in the 90–100 km range. This is consistent with a large body of data that argues for MORB melting to start in the garnet peridotite facies [Salters and Hart, 1989, 1991; Beattie, 1993].

Alternatively, the wide separation between the N-MORB and Ta'u-Ofol arrays could be due to incorrect assignment of mantle olivine Fo values (we used Fo89 for Ta'u-Ofol and Fo90 for the N-MORB). In order to overlap the N-MORB with the lowest temperature Ta'u-Ofol samples, the N-MORB olivine Fo would have to

be increased from Fo90 to Fo93. The N-MORB array would then be subparallel to Tau-Ofu, but 0.5 GPa above it. Conversely, the Fo for the Ta'u-Ofol mantle could be decreased to Fo86, and the array would be subparallel to N-MORB, though 0.5 GPa below it. In either event, to create an overlap would require that the Ta'u-Ofol mantle olivine be 4 Fo units below that of the ambient N-MORB upper mantle. This is simply unsustainable, and further argues for a real P-T difference between N-MORB and the Samoa plume, in either depth of melt origin or melt re-equilibration.

There is one further caveat in understanding these P-T arrays to be trajectories of apparent adiabatic mantle melting—they could also simply be mixing arrays. This process is illustrated in Figure 7c, where an average N-MORB primary melt is mixed with an average of the highest pressure Ofol melts. The result is a slightly curved array that is a reasonable fit to both the N-MORB and Ta'u-Ofol data. This is not to imply an actual mixing of N-MORB and Samoa melts, but simply to show that if a Samoa melt were to stagnate at shallow depth, and re-equilibrate to a low P-T, and then mix with through-going melts derived from great depth, a mixing array would be formed that would explain all of the Samoa data.

10. Re-Equilibration of Migrating Melts

Without further guidance on whether or not melts typically maintain equilibrium during ascent, the best we can say from derived P-T arrays is that the depth and temperature of melt origin is at least as great as the highest "recorded" P and T. However, we can put further constraints on this problem by consideration of the kinetic parameters that govern equilibrium between olivine and silicate melt. Let us then examine the melt transport conditions that are required for a melt to maintain equilibration with the host mantle, or for a melt to re-equilibrate with a mantle wall rock. While Hart [1993] estimated that melts traversing a fractal plumbing system would lose touch with their surroundings over a rather small ascent range (~1–5 km), equilibration length scales may be extended due to a decreasing melting rate as mantle ascent is slowed by impingement on the overlying lithosphere. The simple fractal system considered by Hart [1993] was driven by melting that occurs during mantle upwelling. However, once the melt enters the overlying lithosphere, the transport system and the imposed fluxes will be less simply constrained.

Figure 8 shows the relationship between melt velocities in a transport pipe, and the radius of the pipe, for two closure temperatures defined by the P-T conditions at each end of the P-T array in Figure 7a. We model the diffusive equilibrium of a melt in a cylindrical geometry, equilibrating with a peridotitic wall rock [see Hart, 1993; Jaeger, 1956]. We use the interdiffusion of Mg-Fe in olivine (Fo91) as the rate-limiting step for setting the P-T conditions expressed in the Lee *et al.* [2009] parameterization. Obviously, this is simplistic, as the other mantle phases are also important, and also likely to be diffusively slower (Opx, Cpx, and garnet). On the other hand, the model ignores the certain solution-redeposition that is implied by a melting reaction—and this will tend to speed up the chemical equilibration process. Additionally, the re-equilibration model is insensitive to the total amount of melt that a conduit has carried, as long as the number of mantle phases is maintained, and the bulk composition is held within certain limits.

The closure-temperature formulation from Dodson [1973] is adapted to the boundary conditions and geometry given above (see caption to Figure 8). Two closure-temperature lines are calculated and shown in Figure 8; one for the deepest sample (1600°C) and one for the shallowest sample (1435°C) from the Ofol-Tau array in Figure 7a. Closure temperatures to the NE of each line are higher (nonequilibrated) and to the SW are lower (equilibrated), so the lines demarcate these two domains. A melt plotting in between the lines is thus equilibrated at 1600° but not at 1435°. Melts in a pipe with a large radius and/or high velocity relative to the 1600° line will have frozen in conditions at depths greater than 130 km. Melts in small pipes and/or low velocities relative to the 1435° line are yet to have their P-T conditions frozen in.

For comparison, a re-equilibration curve for the fractal system of Hart [1993] is also shown in Figure 8, and this tends to show an enlarged region of nonequilibration. Note that most choices of model parameters in the fractal model tend to align along a SW-NE field, and this intersection with the closure-temperature arrays may help constrain realistic values for conduit sizes and melt velocities. One constraint on melt velocities is provided by U-series measurements—specifically the fact that ^{226}Ra - ^{230}Th disequilibria is common in many oceanic hot spot systems [Bourdon and Sims, 2003]. This implies a total melt transport plus storage time of less than, say, 5000 years (the half-life of ^{226}Ra is 1600 years). For melting depths of 100 km, melt velocities of greater than 20 m/yr are thus signified. This velocity is curiously close to the demarcation

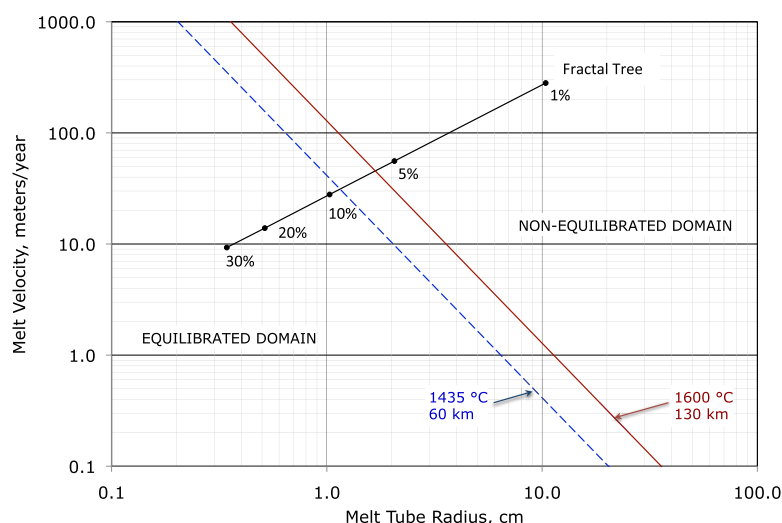


Figure 8. Equilibration Domains in melt radius versus velocity space, defined by Closure Temperatures for the deepest and shallowest samples from Figure 7a. Closure temperatures as defined by Dodson [1973] are for spherical diffusion into an infinite medium, where 83% equilibration occurs at $0.13 (\text{radius})^2/\text{diffusion coefficient}$. In contrast, 83% equilibration times for a melt tube of given radius embedded in olivine is given by $(\text{radius})^2/\text{diffusion coefficient}$ [Hart, 1993; Jaeger, 1956]. Thus to adjust the Dodson formulation to the more realistic geometry as used by Hart [1993], the Dodson times are increased by a factor of 7.7. The cooling rate of Dodson is converted to a melt velocity by using the T-P slope of $2.5^\circ/\text{km}$ for the Ofo-Ta'u array in Figure 7a. The diffusion data are given by Mg-Fe interdiffusion in Fo 91 olivine, with both temperature and pressure dependencies invoked, as published by Chakraborty [1997] and Holzapfel et al. [2007]. Two closure-temperature curves are shown, for both the highest and lowest temperature ends of the Ofo-Ta'u array in Figure 7a (1435°C , 2.0 GPa, $D = 4.7 \times 10^{-16} \text{ m}^2/\text{s}$; 1600°C , 4.25 GPa, $D = 1.3 \times 10^{-15} \text{ m}^2/\text{s}$). The equations of the two curves are: $V = 41.37/r^2$ (dashed) and $V = 128.2/r^2$ (solid). For comparison, an equilibration curve is shown for the fractal tree melt transport model of Hart [1993], using his model C parameters and the updated Fe-Mg interdiffusion data. Points for 1, 5, 10, 20, and 30% equilibration are given. In this case, the melt tube radius increases as the fractal tree branches continue to coalesce. To "lock in" P-T conditions as represented in Figure 7a, melts must have a V-r trajectory that crosses the domain boundary lines from the SW to the NE quadrant, i.e., melts migrating into larger tubes and/or increasing in velocity.

between equilibrated and nonequilibrated domains in Figure 8. There is little else that we know about melt conduit dimensions or melt migration velocities. It would however seem reasonable to expect continuous equilibration and/or re-equilibration events to be common, and it may in fact be difficult to see through to the starting P-T conditions of most melting scenarios.

So what might we conclude from all of this regarding the argument as to whether hot spots (plumes) are hotter than ambient upper (MORB) mantle? Were we to take the maximum potential temperatures indicated for Samoa (1550°C) and primary N-MORB (1360°C), we would conclude (as others have done) that Samoa (and plumes in general) are $\sim 200^\circ\text{C}$ hotter than ambient upper mantle. Given the arguments embodied in Figure 8, however, it could also be argued that N-MORB are as hot as plume mantle, but that virtually all N-MORB have re-equilibrated to relatively low "potential" temperatures due to the particular aspects of their shallow plumbing system (or that some N-MORB re-equilibrate, and the rest are mixed with these, thereby losing their high P-T signature). The geodynamic setting of MORB generation would seem to favor large conduits and rapid ascent velocities, thus making it unlikely that MORBs will have a significant re-equilibration length scale; this argues that the measured P-T values are in fact representative of actual melt generation levels.

11. Summary

1. Alkali basalts from Ta'u volcano and Ofu/Olosega volcano have virtually identical Sr, Nd, Pb, and Hf isotope ratios, indicating long-term evolution in a common mantle source. Based on $^3\text{He}/^4\text{He}$ ratios of up to 33.8 Ra, this mantle source is FOZO in character.

2. Whole rock trace element spidergrams from both volcanoes are strongly LIL-enriched (Ta ~ 70 –90 times BSE). Ofo volcano is $\sim 20\%$ more enriched than Ta'u, and this can be attributed to a smaller degree of melting for Ofo volcano (5.2% versus 6.5% for Ta'u).

3. Mantle source spidergrams constructed for the two volcanoes using melting models show similar patterns, though with statistical differences for U/Nb, K, Sr, Yb, and Lu. Bulk partition coefficients for melting range from ~ 0.001 for Th to ~ 0.10 for Lu.
4. The mantle source spidergram for Ofof has strong positive Nb and Ta anomalies, and a smaller Ti anomaly, as is common in basalts with high $^3\text{He}/^4\text{He}$ and FOZO end-member characteristics.
5. Sm/Nd and Lu/Hf ratios of the sources are low (subchondritic to slightly superchondritic), in contrast to the high (superchondritic) $^{143}\text{Nd}/^{144}\text{Nd}$ and $^{176}\text{Hf}/^{177}\text{Hf}$ isotopic ratios. This dichotomy argues for a recent source enrichment event, possibly autometasomatic in nature.
6. P-T modeling of the basalts defines a tight colinear melting array, with Ofof melts lying at higher temperatures and pressures than Ta'u melts: 1500–1600°C and 90–130 km versus 1430–1560°C and 60–115 km. We argue that this data do not necessarily indicate a range of mantle potential temperatures for these volcanoes, but rather a range of final melt-rock equilibration depths. Diffusion calculations based on Fe-Mg exchange between olivine and melt show that equilibration is likely as long as the melt velocity (in meters per year) is $<40/(\text{conduit radius in cm})^2$.
7. In this “equilibration” interpretation, the highest potential temperature for the Samoan plume would be $\sim 1550^\circ\text{C}$; this is $\sim 200^\circ\text{C}$ higher than the highest observed potential temperature for primary N-MORB. We also show that typical P-T arrays can be well characterized as mixing lines.

Acknowledgments

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